

Neutron stars from skyrmion branes

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- Main aim: Describe baryonic matter on all scales from **finite atomic nuclei** to **dense infinite nuclear matter**

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- Main aim: Describe baryonic matter on all scales from **finite atomic nuclei** to **dense infinite nuclear matter**
- The Skyrme model can be used to model **neutron crystals**, which exist under high pressure inside neutron stars [*Nucl. Phys. B* **262** 133–143 (1985)]

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- Can we obtain a **single EoS** that yields neutron stars with crusts?



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- Can we obtain a **single EoS** that yields neutron stars with crusts?
- Can these neutron stars have sufficient maximal masses?



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- In the large N_c -limit, QCD can be reduced to an effective field theory of mesons
 - Skyrme's idea [*Proc. R. Soc. Lond. A* **260** 127–138 (1961)]: effective mesonic Lagrangian involving only **pions**, with **baryons emerging as stable topological solitons**

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- These are encoded in the Skyrme field

$$\varphi = \begin{pmatrix} \sigma + i\pi_3 & i\pi_1 + \pi_2 \\ i\pi_1 - \pi_2 & \sigma - i\pi_3 \end{pmatrix} \in \text{SU}(2)$$

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$$\mathrm{SU}(2) \ni \begin{pmatrix} \sigma + i\pi_3 & i\pi_1 + \pi_2 \\ i\pi_1 - \pi_2 & \sigma - i\pi_3 \end{pmatrix} \leftrightarrow (\sigma, \pi_1, \pi_2, \pi_3) \in S^3$$

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- $\Rightarrow$ Skyrme field is now a map $\varphi : S^3 \rightarrow \text{SU}(2) \cong S^3$
- Disjoint homotopy classes labelled by $B \in \pi_3(S^3) = \mathbb{Z}$
- $\Rightarrow$ Fields are **topologically stable** and B is identified with the **baryon number**

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Generalized Skyrme Lagrangian

- Skyrme's original model:

$$\mathcal{L}_{24} = \frac{F_\pi^2}{16\hbar} g^{\mu\nu} \text{Tr}(L_\mu L_\nu) + \frac{\hbar}{32e^2} g^{\mu\alpha} g^{\nu\beta} \text{Tr}([L_\mu, L_\nu][L_\alpha, L_\beta])$$
$$\vec{L}_\mu = \varphi^\dagger \partial_\mu \varphi \in \mathfrak{su}(2)$$

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$$\vec{L}_\mu = \vec{\varphi}^\dagger \partial_\mu \vec{\varphi} \in \mathfrak{su}(2)$$

- This is $(\text{SU}(2) \times \text{SU}(2))/\mathbb{Z}_2 \cong \text{SO}(4)$ invariant and the **pions are massless**

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- Boundary condition $\varphi(\vec{x} \rightarrow \infty) = \mathbb{I}_2$ spontaneously breaks chiral $\text{SO}(4)$ symmetry to an isospin $\text{SO}(3)$ symmetry, which acts on $\vec{\pi}$

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- The standard massive Skyrme model includes the **pion mass potential**:

$$\mathcal{L}_{024} = -\frac{F_\pi^2 m_\pi^2}{8\hbar^3} \text{Tr}(\mathbb{I}_2 - \varphi) + \frac{F_\pi^2}{16\hbar} g^{\mu\nu} \text{Tr}(L_\mu L_\nu) + \frac{\hbar}{32e^2} g^{\mu\alpha} g^{\nu\beta} \text{Tr}([L_\mu, L_\nu][L_\alpha, L_\beta])$$

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Generalized Skyrme Lagrangian

- Generalized Skyrme model includes a sextic term [*Phys. Lett. B* **154**, 101–106 (1985)], which is related to the ω -Skyrme model [*Phys. Lett. B* **137**, 251–256 (1984)]:

$$\begin{aligned}\mathcal{L}_{0246} = & -\frac{F_\pi^2 m_\pi^2}{8\hbar^3} \text{Tr}(\mathbb{I}_2 - \varphi) + \frac{F_\pi^2}{16\hbar} g^{\mu\nu} \text{Tr}(L_\mu L_\nu) + \frac{\hbar}{32e^2} g^{\mu\alpha} g^{\nu\beta} \text{Tr}([L_\mu, L_\nu][L_\alpha, L_\beta]) \\ & - \pi^4 \lambda^2 g^{\mu\nu} \mathcal{B}_\mu \mathcal{B}_\nu, \quad \lambda^2 = g_\omega^2 / (2\pi^4 m_\omega^2)\end{aligned}$$

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- Exhibits short range ω -meson-like repulsion while still describing scalar meson effects

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- Baryon d.o.f. not explicitly visible $\rightarrow$ topology: Homotopy invariant $\leftrightarrow$ Baryon number

$$\pi_3(\text{SU}(2)) = \mathbb{Z} \ni B = \int_{\mathbb{R}^3} d^3x \sqrt{-g} \mathcal{B}^0, \quad \mathcal{B}^\mu = \frac{1}{24\pi^2 \sqrt{-g}} \epsilon^{\mu\nu\rho\sigma} \text{Tr}(L_\nu L_\rho L_\sigma)$$

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- Baryons realized as non-perturbative excitations of the pions $\Rightarrow$ solutions of the Euler–Lagrange field equations - topological solitons (**skyrmions**)

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- In Skyrme units the energy-momentum tensor is

$$T_{\mu\nu} = -\frac{2}{\sqrt{-g}} \frac{\partial(\sqrt{-g}\mathcal{L}_{0246})}{\partial g^{\mu\nu}} \quad \frac{\pi^4 \lambda^2 e^4 F_\pi^2}{2\hbar^3} = c_6 \boxtimes$$

$$= -\text{Tr}(L_\mu L_\nu) - \frac{1}{4} g^{\alpha\beta} \text{Tr}([L_\mu, L_\alpha][L_\nu, L_\beta]) + 2c_6 \mathcal{B}_\mu \mathcal{B}_\nu + g_{\mu\nu} \mathcal{L}_{0246}$$

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$$T_{\mu\nu} = -\text{Tr}(L_\mu L_\nu) - \frac{1}{4}g^{\alpha\beta} \text{Tr}([L_\mu, L_\alpha][L_\nu, L_\beta]) + 2c_6 \mathcal{B}_\mu \mathcal{B}_\nu + g_{\mu\nu} \mathcal{L}_{0246}$$

- The adimensional static energy is thus ($T_{00} = \mathcal{E}_{\text{stat}} + \mathcal{E}_{\text{kin}}$)

$$\begin{aligned} M_B(\varphi, g) &= \int_{\mathbb{R}^3} d^3x \sqrt{-g} \mathcal{E}_{\text{stat}} \\ &= \int_M d^3x \sqrt{-g} \left\{ -\frac{1}{2} g^{ij} \text{Tr}(L_i L_j) - \frac{1}{16} g^{ik} g^{jl} \text{Tr}([L_i, L_j][L_k, L_l]) \right. \\ m = \frac{2m_\pi}{F_\pi e} &\rightarrow \left. + m^2 \text{Tr}(\mathbb{I}_2 - \varphi) + c_6 \frac{\epsilon^{ijk} \epsilon^{abc}}{(24\pi^2 \sqrt{-g})^2} \text{Tr}(L_i L_j L_k) \text{Tr}(L_a L_b L_c) \right\} \end{aligned}$$

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Generalized Skyrme model

- We are interested in **static** solutions and adopt the usual Skyrme units of length $\tilde{L} = 2\hbar/eF_\pi$ and energy $\tilde{E} = F_\pi/4e$
- In Skyrme units the energy-momentum tensor is

$$T_{\mu\nu} = -\text{Tr}(L_\mu L_\nu) - \frac{1}{4}g^{\alpha\beta}\text{Tr}([L_\mu, L_\alpha][L_\nu, L_\beta]) + 2c_6\mathcal{B}_\mu\mathcal{B}_\nu + g_{\mu\nu}\mathcal{L}_{0246}$$

- The adimensional static energy is thus

$$M_B(\varphi, g) = \int_{\mathbb{R}^3} d^3x \sqrt{-g} \left\{ -\frac{1}{2}g^{ij}\text{Tr}(L_i L_j) - \frac{1}{16}g^{ik}g^{jl}\text{Tr}([L_i, L_j][L_k, L_l]) \right. \\ \left. + m^2\text{Tr}(\mathbb{I}_2 - \varphi) + c_6 \frac{\epsilon^{ijk}\epsilon^{abc}}{(24\pi^2\sqrt{-g})^2}\text{Tr}(L_i L_j L_k)\text{Tr}(L_a L_b L_c) \right\}$$

- Skyrmions are **energy minimizing** static solutions of the Euler–Lagrange equations associated to M_B

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Linking in the Skyrme model

Baryon number as the linking number of vortices



- The Skyrme field can be written in terms of two vortices $\psi_1, \psi_2 \in \mathbb{C}$ as

$$\phi = \begin{pmatrix} \psi_1 & -\bar{\psi}_2 \\ \psi_2 & \bar{\psi}_1 \end{pmatrix}$$

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- Consider the Hopf map $H : S^3 \rightarrow S^2$ due to the Hopf fibration $S^1 \hookrightarrow S^3 \rightarrow S^2$

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- The map $\varphi : \mathbb{R}^3 \cup \{\infty\} \cong S^3 \rightarrow S^3$ of degree B has Hopf charge $Q = B$ under the Hopf map $H : S^3 \rightarrow S^2$ [*Phys. Rev. D* **101**, 065011 (2020)]

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- Distinct regular points on S^2 under $H \circ \varphi : \mathbb{R}^3 \cup \{\infty\} \rightarrow S^2$ have preimages on $\mathbb{R}^3 \cup \{\infty\}$ that are linked $Q = B$ times
- Antipodal points on S^2 identified as **vortex lines** (zeros) and are linked $Q = B$ times

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$B = 1$  skyrmion [*Proc. R. Soc. Lond. A* **260** 127-138 (1961)]

- The hedgehog field is $\varphi(\vec{x}) = \exp(if(r)\vec{x} \cdot \vec{\tau})$

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$$B = -\frac{1}{2\pi^2} \int_0^\infty \frac{\sin^2 f}{r^2} \frac{df}{dr} 4\pi r^2 dr = \frac{1}{\pi} f(0) = 1$$

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- For the hedgehog ansatz, the (massless) static energy is

$$M_1 = 4\pi \int_0^\infty \left[r^2 \left(\frac{df}{dr} \right)^2 + 2 \sin^2 f \left(1 + \left(\frac{df}{dr} \right)^2 \right) + \frac{\sin^4 f}{r^2} \right] dr$$

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- E-L equations reduce to a 2nd order non-linear ODE that can only be solved numerically,

$$\left(r^2 + 2 \sin^2 f \right) \frac{d^2 f}{dr^2} + 2r \frac{df}{dr} + \sin 2f \left[\left(\frac{df}{dr} \right)^2 - 1 - \frac{\sin^2 f}{r^2} \right] = 0$$

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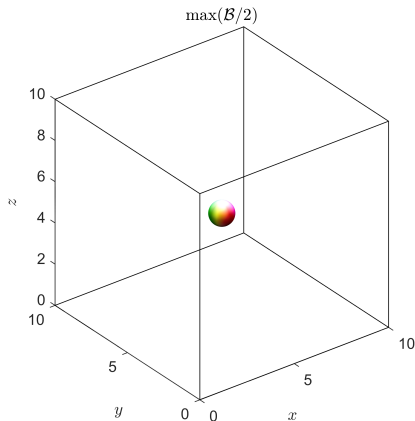
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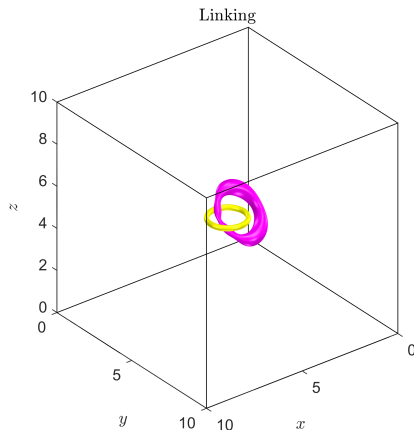
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(a) Isosurface plot of the baryon density $\mathcal{B}_0$



(b) Linking of two preimages

How to construct larger Skyrmions



- Asymptotic interactions of two $B = 1$ skyrmions have preferred orientation (**attractive channel**) [*Commun. Math. Phys.* **245** 123–147 (2004)]

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- One could also relate Skyrmions to **instantons** via the Atiyah–Manton approximation [*Phys. Lett. B* **222** 438–442 (1989)] or using ADHM data [*Nonlinearity* **35**, 3944–3990 (2022)]
- Quite large skyrmions (up to $B = 108$) have been constructed by gluing α -particles together [*Proc. R. Soc. A* **463** 261–279 (2007)] or using a multi-layer RMA based on the **Skyrme crystal** [*Phys. Rev. D* **87** 0850834 (2013)]

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$$\varphi(r, z) = \exp \left[\frac{if(r)}{1 + |R|^2} \begin{pmatrix} 1 - |R|^2 & 2\bar{R} \\ 2R & |R|^2 - 1 \end{pmatrix} \right],$$

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- Rational maps for $B = 1, \dots, 4$:

B	$R(z)$	Symmetry
1	z	$O(3)$
2	z^2	$O(2) \times \mathbb{Z}$
3	$\frac{z^3 - \sqrt{3}iz}{\sqrt{3}iz^2 - 1}$	T_d
4	$\frac{z^4 + 2\sqrt{3}iz^2 + 1}{z^4 - 2\sqrt{3}iz^2 + 1}$	O_h

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- Substituting the RMA into the massless static energy functional yields

$$M_B = 4\pi \int_0^\infty r^2 \left\{ \left(\frac{df}{dr} \right)^2 + 2B \sin^2 f \left(\left(\frac{df}{dr} \right)^2 + 1 \right) + \mathcal{F} \frac{\sin^4 f}{r^2} \right\} dr,$$

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- $\mathcal{F}$ is the purely angular integral to be minimised for choice of rational map R :

$$\mathcal{F} = \frac{1}{4\pi} \int \left(\frac{1 + |z|^2}{1 + |R|^2} \left| \frac{dR}{dz} \right| \right)^4 \frac{2i dz d\bar{z}}{(1 + |z|^2)^2}.$$

Skyrmions from rational maps [*Nucl. Phys. B* **510**, 507–537 (1998)]



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- Substituting the RMA into the massless static energy functional yields

$$M_B = 4\pi \int_0^\infty r^2 \left\{ \left(\frac{df}{dr} \right)^2 + 2B \sin^2 f \left(\left(\frac{df}{dr} \right)^2 + 1 \right) + \mathcal{F} \frac{\sin^4 f}{r^2} \right\} dr,$$

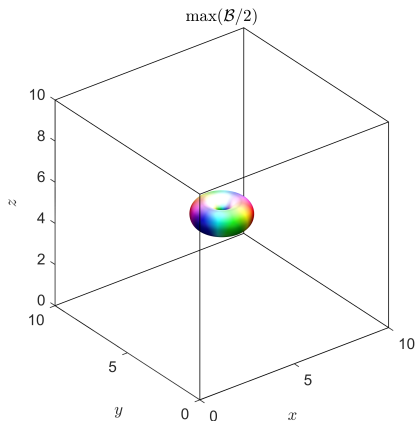
- $\mathcal{F}$ is the purely angular integral to be minimised for choice of rational map R :

$$\mathcal{F} = \frac{1}{4\pi} \int \left(\frac{1 + |z|^2}{1 + |R|^2} \left| \frac{dR}{dz} \right| \right)^4 \frac{2i dz d\bar{z}}{(1 + |z|^2)^2}.$$

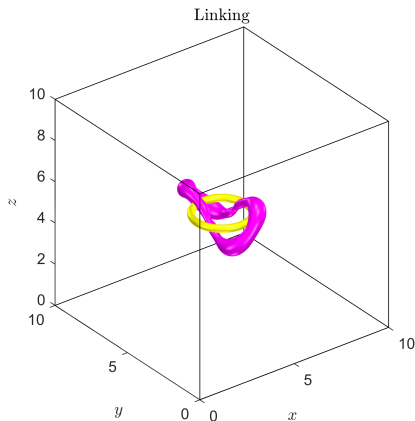
- Optimising $\mathcal{F}$ and the profile function $f(r)$ gives approximate Skyrmions, but further numerical relaxation is required to find true Skyrmions.



$B = 2$ skyrmion [*Phys. Lett. B* **195**, 235–239 (1987)]



(a) Isosurface plot of the baryon density $\mathcal{B}_0$



(b) Linking of two preimages

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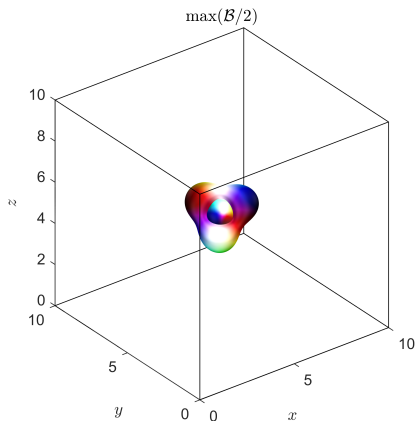
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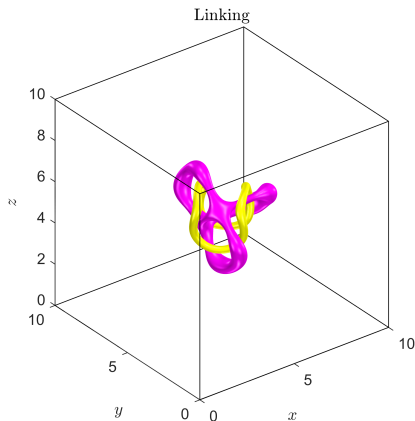
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$B = 3$ skyrmion [*Phys. Lett. B* 235, 147–152 (1990)]



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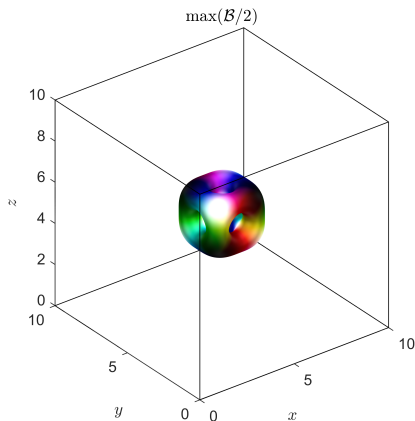
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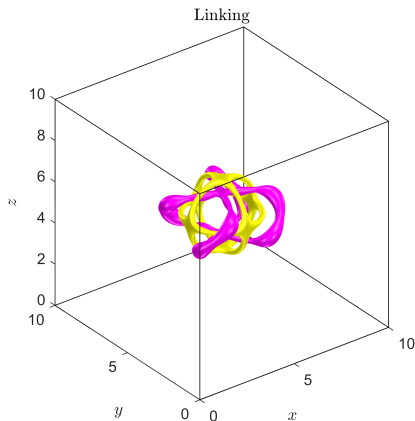
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$B = 4$ skyrmion [*Phys. Lett. B* 235, 147–152 (1990)]



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Motivation of Skyrme crystals



- Aim: construct an **equation of state** (EoS) to model neutron stars

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Motivation of Skyrme crystals

- Aim: construct an **equation of state** (EoS) to model neutron stars
- ⇒ We need to understand **phases** and **phase transitions** of nuclear matter

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Motivation of Skyrme crystals

- Aim: construct an **equation of state** (EoS) to model neutron stars
- ⇒ We need to understand **phases** and **phase transitions** of nuclear matter
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Motivation of Skyrme crystals

- Aim: construct an **equation of state** (EoS) to model neutron stars
- ⇒ We need to understand **phases** and **phase transitions** of nuclear matter
- Ground state of dense nuclear matter has a **crystalline** structure in the classical approximation
 - In order to determine skyrmion crystals, we first need to define what a crystal really is!

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Skyrmion crystals



- Skyrme crystals are energy minimizing maps

$$\varphi : \mathbb{R}^3 / \Lambda_\circ \rightarrow \mathrm{SU}(2), \quad \Lambda_\circ = \{n_1 \vec{X}_1 + n_2 \vec{X}_2 + n_3 \vec{X}_3 : n_i \in \mathbb{Z}\}$$

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- Key idea [*Comm. Math. Phys.* **332** 355-377 (2014)]: Identify all 3-tori via diffeomorphism (with $\mathbb{T}^3 \equiv \mathbb{R}^3 / \mathbb{Z}^3$)

$$F : (\mathbb{T}^3, g) \rightarrow (\mathbb{R}^3 / \Lambda, d), \quad (x^1, x^2, x^3) \mapsto x^1 \vec{X}_1 + x^2 \vec{X}_2 + x^3 \vec{X}_3$$

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- Vary metric g_i with $g_0 = F^* d \iff$ vary lattice Λ_s with $\Lambda_0 = \Lambda$

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- Vary metric g with $g_0 = F^* d \iff$ vary lattice Λ_s with $\Lambda_0 = \Lambda$
- Energy minimized over variations of $g \iff$ energy minimizing period lattice $\Lambda_\circ$.

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Summary of [*J. Math. Phys.* **64** 103503 (2023)]

- For fixed $\mathcal{L}_{024}$ -field φ , there always **exists** a critical point of $M_B(\varphi, g)$ w.r.t. variations of g and it is in fact a **unique** c.p. (generalizes to $\mathcal{L}_{0246}$ -model)

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- Four crystal solutions were found for unit cells with charge $B_{\text{cell}} = 4$
- These are the φ_{FCC} , φ_α , φ_{string} and φ_{brane} crystals
- The φ_{FCC} -crystal [Phys. Lett. B 208 491–494 (1988)] can be obtained from a Fourier series-like expansion as an initial configuration [Nucl. Phys. A 501 801–812 (1989)],

$$\sigma = -c_1 c_2 c_3, \quad \pi_1 = s_1 \sqrt{1 - \frac{s_2^2}{2} - \frac{s_3^2}{2} + \frac{s_2^2 s_3^2}{3}}, \quad \text{and cyclic,}$$

where $s_i = \sin(2\pi x^i / L)$ and $c_i = \cos(2\pi x^i / L)$, with initial metric $g = L^2 \mathbb{I}_3$.

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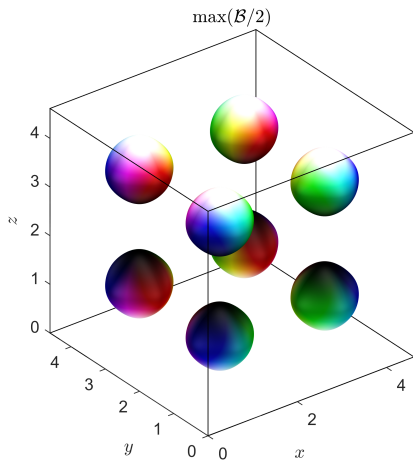
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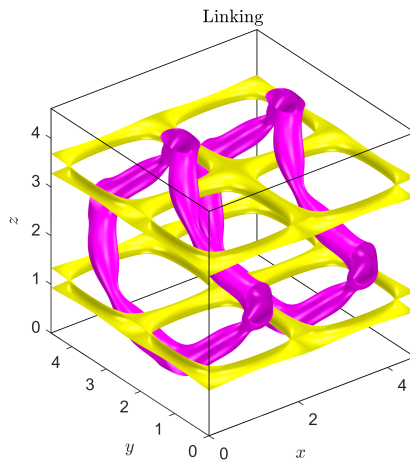
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Half-skyrmion (FCC) crystal [*Phys. Lett. B* **208** 491–494 (1988)]



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Summary of [*J. Math. Phys.* **64** 103503 (2023)]

- From φ_{FCC} , the other three crystals can be constructed by applying a chiral $\text{SO}(4)$ transformation $Q \in \text{SO}(4)$, such that $\varphi = Q\varphi_{\text{FCC}}$, and minimizing $\mathcal{M}_B$ w.r.t. variations of φ and g

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- ⇒ Should yield a **lower compression modulus** than previous studies

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 - $\Delta E = E_{\text{isolated}} - E_{\text{min}}$ is minimized for the choice of crystal φ_{brane}
- ⇒ Should yield a **lower compression modulus** than previous studies
- ⇒ **Brane crystal** is an ideal candidate for **dense nuclear matter**

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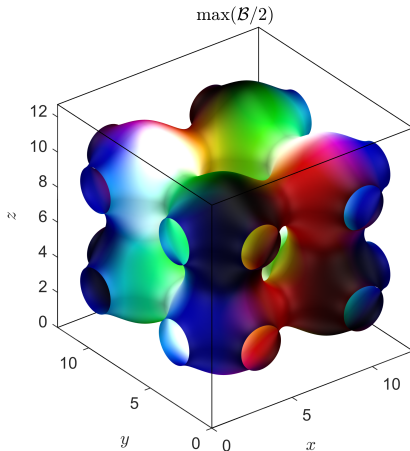
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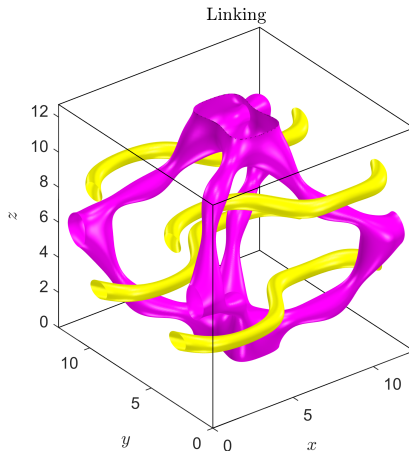
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Brane or domain wall crystal [*J. Math. Phys.* **64** 103503 (2023)]



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Varying the metric on $\mathbb{T}^3$

- Let g be a smooth one-parameter family of metrics on $\mathbb{T}^3$ with $g_0 = F^* d$



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- Let g_s be a smooth one-parameter family of metrics on $\mathbb{T}^3$ with $g_0 = F^*d$
- Set $\delta g = \partial_s g_s|_{s=0} \in \Gamma(\odot^2 T^* \mathbb{T}^3)$ (symmetric 2-covariant tensor field on $\mathbb{T}^3$)

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- Inner product on the space of 2-covariant tensor fields $\langle A, B \rangle_g = A_{ij} g^{jk} B_{kl} g^{li}$

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- First variation of M_B w.r.t. g_s is

$$\left. \frac{dM_B(\varphi, g_s)}{ds} \right|_{s=0} = \int_{\mathbb{T}^3} d^3x \sqrt{g} \langle S(\varphi, g), \delta g \rangle_g, \quad S(\varphi, g) \in \Gamma(\odot^2 T^* \mathbb{T}^3)$$

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- $S(\varphi, g)$ is the **stress-energy tensor**:

$$S_{ij} = \frac{1}{2} \left[m^2 \text{Tr}(\text{Id} - \varphi) - \frac{1}{2} g^{kl} \text{Tr}(L_k L_l) - \frac{1}{16} g^{km} g^{ln} \text{Tr}([L_k, L_l][L_m, L_n]) - c_6 (B_0)^2 \right] g_{ij} \\ + \frac{1}{2} \text{Tr}(L_i L_j) + \frac{1}{8} g^{kl} \text{Tr}([L_i, L_k][L_j, L_l]).$$

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Numerical minimization of the field and lattice



- Fix $\varphi : \mathbb{T}^3 \rightarrow SU(2)$ and think of the energy as a map $E_\varphi : SPD_3 \rightarrow \mathbb{R}$ such that $E_\varphi := M_B(\varphi|_{\text{fixed}}, g)$

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Numerical minimization of the field and lattice

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- We use arrested Newton flow on SPD_3 to minimize E_φ w.r.t. g

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- We use arrested Newton flow on SPD_3 to minimize E_φ w.r.t. g
- Explicitly, we are solving the system of 2nd order ODEs

$$\left. \frac{d^2}{ds^2} \right|_{s=0} (g_{ij})_s = -\frac{\partial E_\varphi}{\partial g_{ij}} = -\int_{\mathbb{T}^3} d^3x \sqrt{g} S_\varphi^{ij}, \quad (g_{ij})_0 = \vec{X}_i \cdot \vec{X}_j$$

where $S_\varphi \equiv S(\varphi|_{\text{fixed}}, g)$ is the fixed field stress tensor

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- In conjunction, we minimize $M_B(\varphi, g|_{\text{fixed}})$ w.r.t. φ for some initial field φ_0

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- ⇒ Laddering of minimizations

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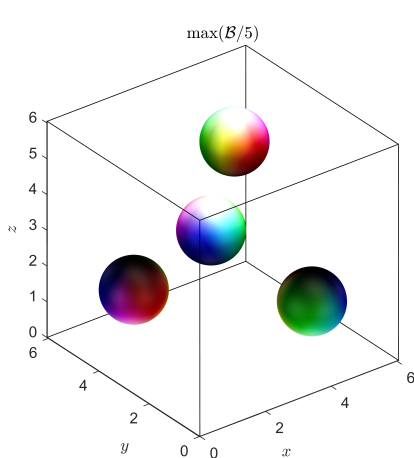
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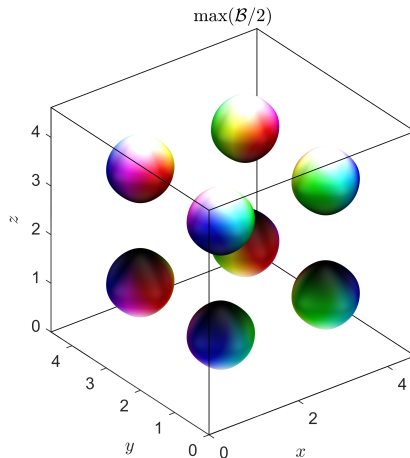
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An example: the FCC to half-skyrmion crystal



(a) Initial configuration of four $B = 1$ hedgehogs arranged in the attractive channel on a cubic grid



(b) Relaxed final solution of the cubic crystal of half-skyrmions

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Phases of skyrmion matter

- Consider fixed baryon density n_B variations of $\mathcal{M}_B(\phi, g)$ w.r.t. g



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$$\left. \frac{d}{ds} \right|_{s=0} \int_{T^3} d^3x \sqrt{g_s} = \frac{1}{2} \int_{T^3} d^3x \sqrt{g} g^{ij} \delta g_{ij} = 0$$

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- Leads to modifying the (fixed φ field) stress-energy tensor via the mapping

$$S_\varphi \mapsto \tilde{S}_\varphi = S_\varphi - \frac{1}{3} \text{Tr}_g(S_\varphi) g$$

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- Convergence criterion becomes $\max(\tilde{S}_\varphi) < \text{tol}$
- This process enables us to determine an **energy-density** curve
- This is key to obtaining an **equation of state** within our framework

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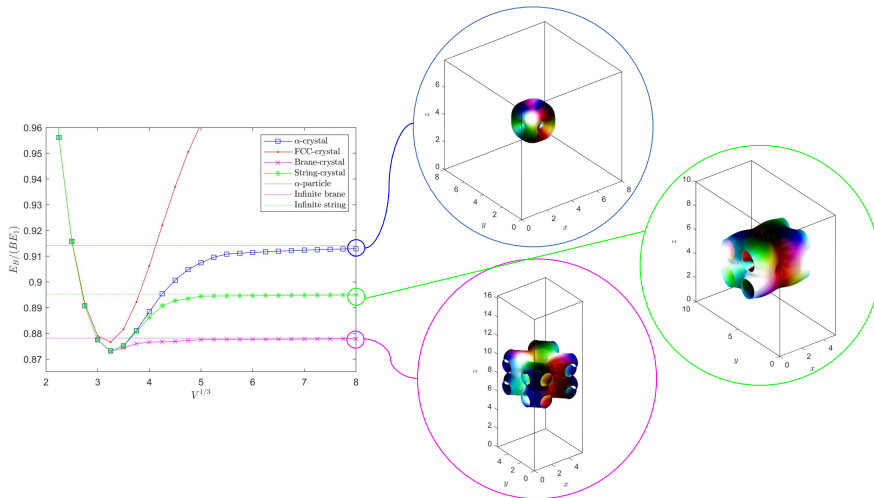
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Isospin quantization



- Non-renormalizable theory $\Rightarrow$ isospin asymmetry is included by semi-classically quantizing isospin collective coordinates: $\varphi(x) \mapsto \hat{\varphi}(x, t) = A(t)\varphi(x)A^\dagger(t)$ [*Nucl. Phys. B* **262** 133–143 (1985)]

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- Can use a mean-field approximation of a large chunk ($B = N_{\text{cell}}B_{\text{cell}}$) in a generic quantum state with fixed eigenvalue [*Phys. Rev. D* **106** 114031 (2022)]

$$I_3 = \frac{(Z - N)}{2} = -\frac{(1 - 2\gamma_p)}{2}N_{\text{cell}}B_{\text{cell}}, \quad \gamma_p \text{ is the proton fraction}$$

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- $I = I_3$ minimizes the isospin energy since by definition $I^2 \geq I_3^2$
- The isospin correction (per unit cell) to the energy of the crystal is found to be $E_{\text{iso}} = \frac{\hbar^2}{8U_{33}}B_{\text{cell}}^2\delta^2$

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Symmetry energy

- The asymmetry of matter is determined by the isospin asymmetry parameter

$$\delta = (N - Z)/(N + Z) = 1 - 2\gamma_p$$



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Symmetry energy

- The asymmetry of matter is determined by the isospin asymmetry parameter $\delta = (N - Z)/(N + Z) = 1 - 2\gamma_p$
- Binding energy per baryon number of asymmetric nuclear matter is given by

$$\frac{E}{B}(n_B, \delta) = E_N(n_B) + S_N(n_B)\delta^2 + \mathcal{O}(\delta^3), \quad \begin{aligned} n_0 &= 0.160 \text{ fm}^{-3} \\ E_N(n_0) &= 923 \text{ MeV} \\ S_N(n_0) &\approx 30 \text{ MeV} \end{aligned}$$

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- The isospin symmetric binding energy is defined by $E_N = M_B/B$
- The symmetry energy S_N dictates how the binding energy changes going from symmetric ($\delta = 0$) to asymmetric ($\delta \neq 0$) nuclear matter

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- The symmetry energy S_N dictates how the binding energy changes going from symmetric ($\delta = 0$) to asymmetric ($\delta \neq 0$) nuclear matter
- It is obtained from the quantum isospin energy $S_N(n_B) = \frac{E_{\text{iso}}}{B_{\text{cell}}\delta^2} = \frac{\hbar^2}{8U_{33}}V_{\text{cell}}n_B$

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- The isospin symmetric binding energy is defined by $E_N = M_B/B$
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- It is obtained from the quantum isospin energy $S_N(n_B) = \frac{E_{\text{iso}}}{B_{\text{cell}}\delta^2} = \frac{\hbar^2}{8U_{33}}V_{\text{cell}}n_B$
- In [arXiv:2306.04533 (2023)], at saturation we find

$$n_0 = 0.160 \text{ fm}^{-3}, E_N(n_0) = 912 \text{ MeV} \text{ and } S_N(n_0) = 22.7 \text{ MeV}$$

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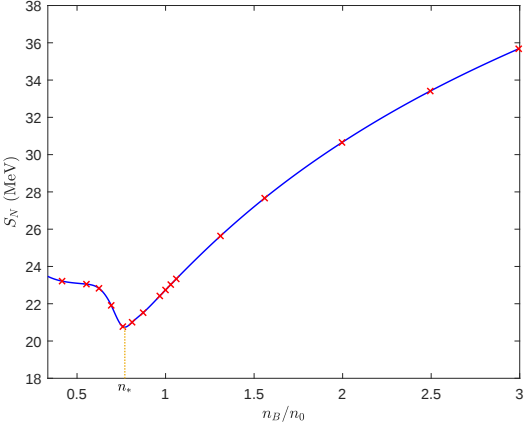
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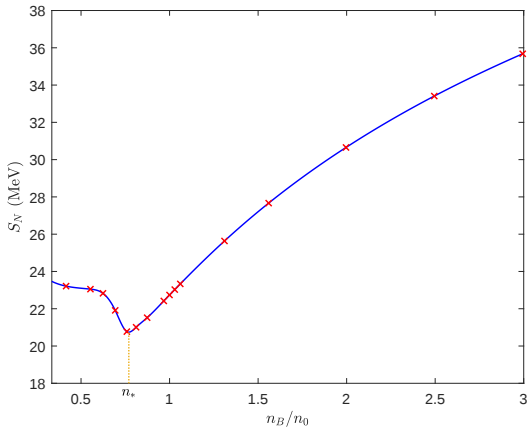
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- **Cusp** below saturation at $n_* \sim 3n_0/4$

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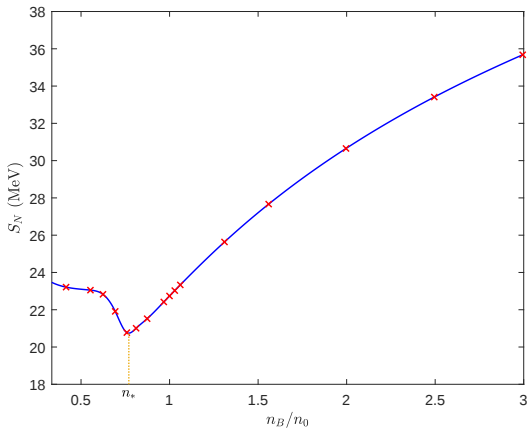
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- **Cusp** below saturation at $n_* \sim 3n_0/4$
- Symmetry energy at zero density $S_N(0) = 23.77$ MeV (finite symmetric nucl. mat.)

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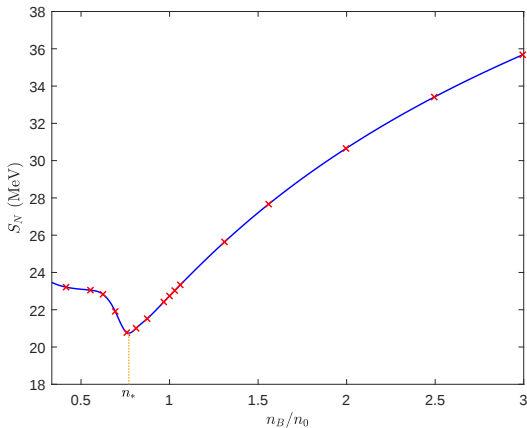
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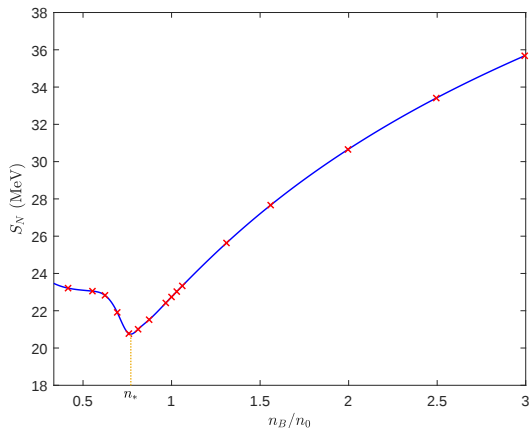
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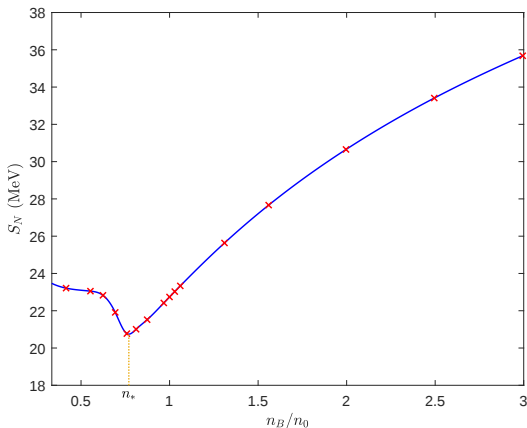
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- Can identify $S_N(0) \sim a_A = 23.7 \text{ MeV}$
- Cusp origin: **phase transition** between **infinite isospin asymmetric nuclear matter** and somewhat **isolated finite nuclear matter**
[arXiv:2306.04533 (2023)]

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Particle fractions of $npe\mu$ matter in β -equilibrium

- Global **charge neutrality** by including background of charged leptons $n_p = n_e + n_\mu$

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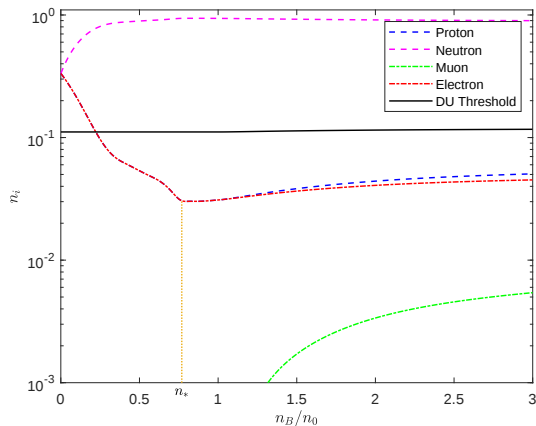
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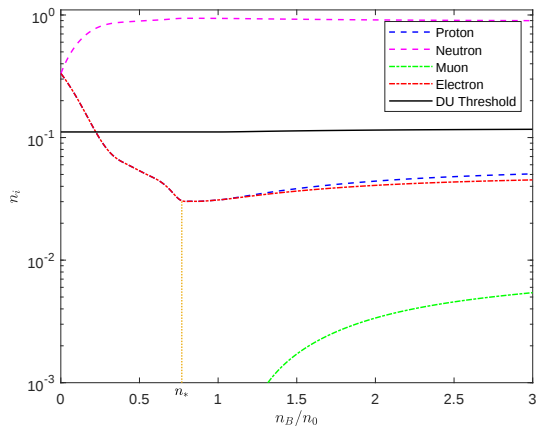
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- Cusp also present at n_*

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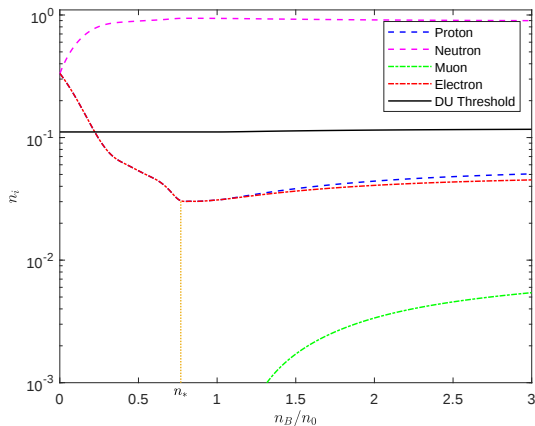
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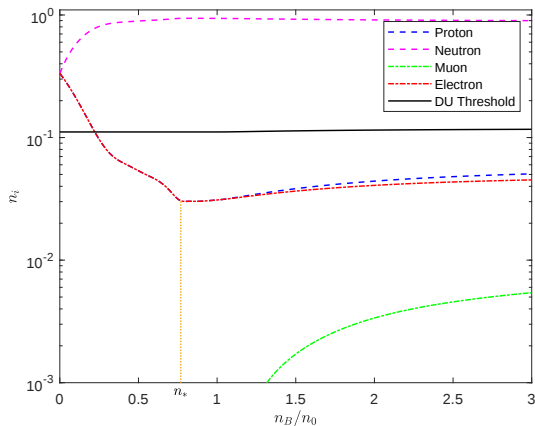
Particle fractions of $npe\mu$ matter in β -equilibrium



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Particle fractions of $npe\mu$ matter in β -equilibrium



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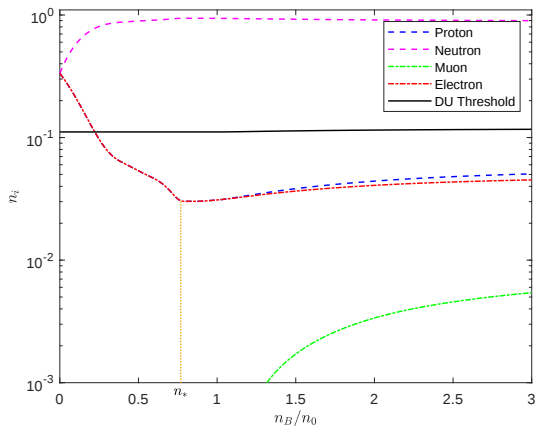
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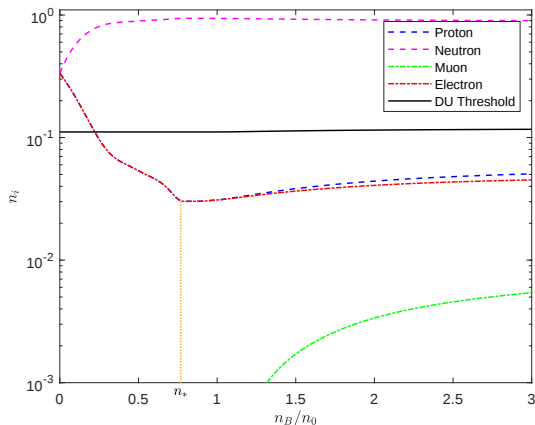
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Particle fractions of $npe\mu$ matter in β -equilibrium



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- The crust of NS is iron rich with $\gamma_p = 0.46$ for ^{56}Fe
- We find as $n_B \rightarrow 0$ then $\gamma_p = 0.5$

⇒ These correspond quite well



Particle fractions of $npe\mu$ matter in β -equilibrium

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 - As n_B increases then so too does n_p and $n_e \rightarrow \mu_e \geq m_\mu = 105.66 \text{ MeV}$
- ⇒ Energetically favourable for muons to appear
- The simultaneous β -decay and electron capture processes allow the calculation of the proton fraction γ_p at a prescribed density n_B [*Phys. Rev. D* **106** 114031 (2022)]
 - Energy of a relativistic Fermi gas at zero temperature (lepton energy)

$$E_l(n_B) = \frac{B_{\text{cell}}}{n_B \hbar^3 \pi^2} \int_0^{k_F} k^2 \sqrt{k^2 + m_l^2} dk, \quad k_F = (3\pi^2 n_l)^{1/3}, \quad n_l = \gamma_l n_B$$

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- Energy per unit cell of β -equilibrated matter

$$E_{\text{cell}}(n_B) = M_B(n_B) + E_{\text{iso}}(n_B) + E_e(n_B) + E_\mu(n_B)$$

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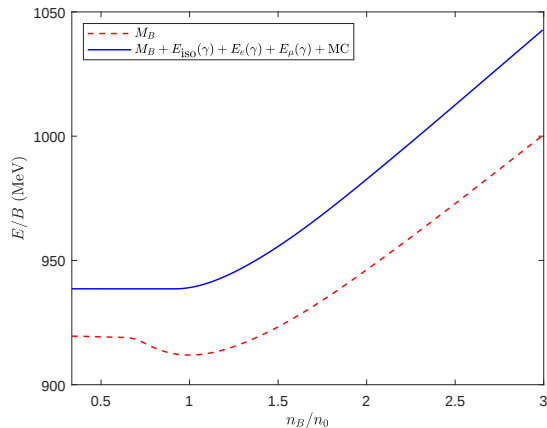
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Isospin asymmetric equation of state



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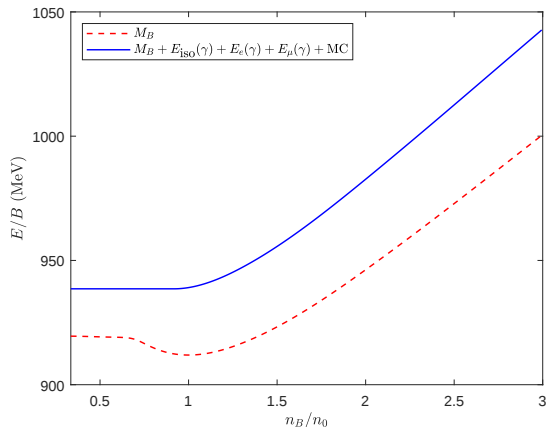
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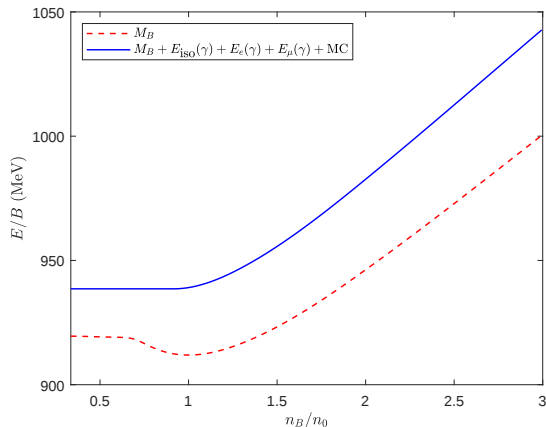


- Can obtain the pressure p and energy density ρ from the $E(n_B)$ curve, with

$$\rho = \frac{E}{V} = \frac{n_B}{B} E_{\text{cell}}$$

$$p = - \frac{\partial E}{\partial V} = \frac{n_B^2}{B} \frac{\partial E_{\text{cell}}}{\partial n_B}$$

Isospin asymmetric equation of state



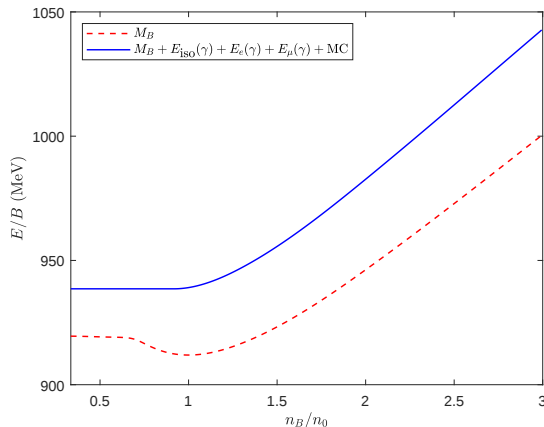
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$\Rightarrow$ Isospin asymmetric nuclear matter
EoS $\rho_{\text{brane}} = \rho_{\text{brane}}(p)$

Isospin asymmetric equation of state



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- $\Rightarrow$ Isospin asymmetric nuclear matter
EoS $\rho_{\text{brane}} = \rho_{\text{brane}}(p)$
- We will use this EoS to obtain NS within the Skyrme model



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Coupling to gravity

- In order to describe neutrons stars within the Skyrme framework, we need to couple the generalized Skyrme model to gravity



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Coupling to gravity

- In order to describe neutrons stars within the Skyrme framework, we need to couple the generalized Skyrme model to gravity
- Introduce the Einstein–Hilbert–Skyrme action

$$S = \frac{1}{16\pi G} \int_{\Sigma} d^4x \sqrt{-g} R + S_{\text{matter}}$$

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- S_{matter} describes matter inside NS
- NS Interior well described by **perfect fluid** of nearly free neutrons & degenerate gas of electrons:

$$T_{\mu\nu} = -\frac{2}{\sqrt{-g}} \frac{\delta S_{\text{matter}}}{\delta g^{\mu\nu}} = (\rho(p) + p)u_{\mu}u_{\nu} + pg_{\mu\nu}$$

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- The energy density ρ and the pressure p are related by the (Brane) crystal EoS
 $\rho(p) = \rho_{\text{brane}}(p)$ [*Phys. Lett. B* **811** 135928 (2020)]

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The Tolman–Oppenheimer–Volkoff system

- Our aim is to calculate $M_{\max}$ and $R_{\max}$ for a NS described by our system

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The Tolman–Oppenheimer–Volkoff system

- Our aim is to calculate $M_{\max}$ and $R_{\max}$ for a NS described by our system
- Need to solve the Einstein equations $G_{\mu\nu} = 8\pi GT_{\mu\nu}$ for some particular choice of $g_{\mu\nu}$

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- Spherically symmetric ansatz of the spacetime metric

$$ds^2 = -A(r)dt^2 + B(r)dr^2 + r^2 (d\theta^2 + \sin^2 \theta d\phi^2) = g_{\mu\nu} dx^\mu dx^\nu$$

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$$ds^2 = -A(r)dt^2 + B(r)dr^2 + r^2 (d\theta^2 + \sin^2 \theta d\phi^2) = g_{\mu\nu} dx^\mu dx^\nu$$

- Substituting this into the Einstein equations $G_{\mu\nu} = R_{\mu\nu} - \frac{1}{2}Rg_{\mu\nu} = 8\pi GT_{\mu\nu}$ yields the TOV system

$$\begin{aligned}\frac{dA}{dr} &= A(r)r \left(8\pi GB(r)p(r) - \frac{1 - B(r)}{r^2} \right) \\ \frac{dB}{dr} &= B(r)r \left(8\pi GB(r)\rho(p(r)) + \frac{1 - B(r)}{r^2} \right) \\ \frac{dp}{dr} &= - \frac{p(r) + \rho(p(r))}{2A(r)} \frac{dA}{dr}\end{aligned}$$

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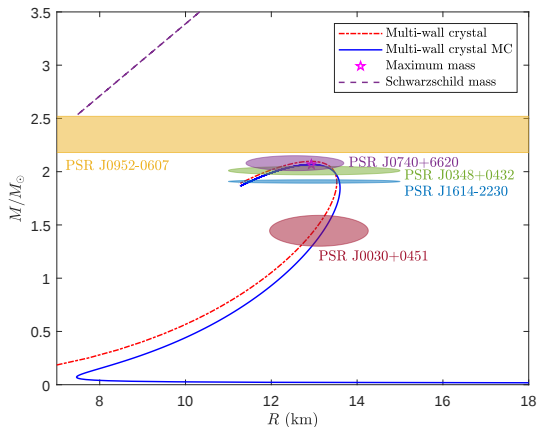
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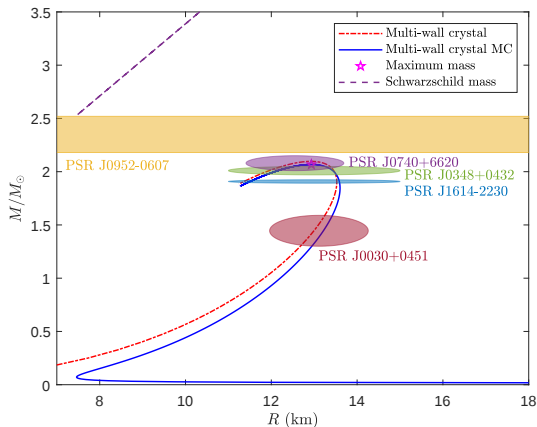
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Neutron star properties and the mass-radius curve



- Mass M obtained from Schwarzschild metric definition outside the star

$$B(R_{\text{NS}}) = \frac{1}{1 - \frac{2MG}{R_{\text{NS}}}}$$

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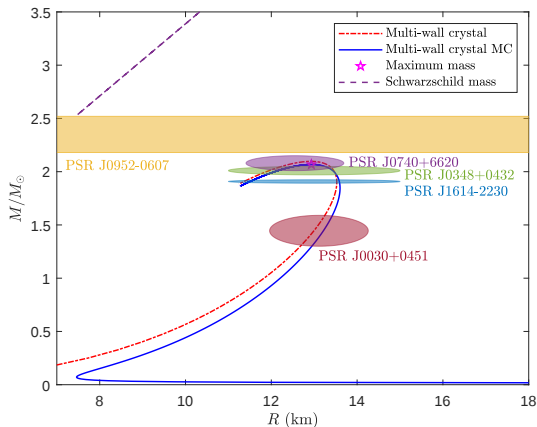
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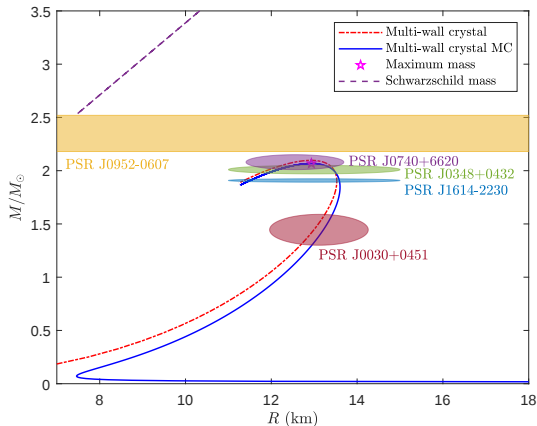


- Mass M obtained from Schwarzschild metric definition outside the star

$$B(R_{\text{NS}}) = \frac{1}{1 - \frac{2MG}{R_{\text{NS}}}}$$

- $M_{\text{max}} = 2.0971M_{\odot}$, occurring for a neutron star of radius $R_{\text{NS}} = 13.12 \text{ km}$.

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$$B(R_{\text{NS}}) = \frac{1}{1 - \frac{2MG}{R_{\text{NS}}}}$$

- $M_{\text{max}} = 2.0971M_{\odot}$, occurring for a neutron star of radius $R_{\text{NS}} = 13.12 \text{ km}$.
- ⇒ Resulting neutron stars agree well with recent NICER/LIGO observational data



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- Cusp structure in the symmetry energy observed in the hidden-local-symmetric (HLS) Skyrme model [*Phys. Rev. C* **83** 025206 (2011)]

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- There is a topological phase transition where the FCC lattice of hedgehog skyrmions fractionalize into half-skyrmions (FCC crystal)
- Analogous to “pseudo-gap” phenomenon in condensed matter physics

Open problems

- Brane solution improves on compressibility at saturation



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- Estimation of the other **SEMF coefficients** a_V, a_S, a_C
 - What does the cusp in our results mean phenomenologically?